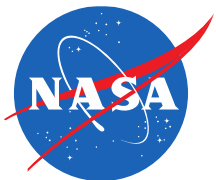


# ***Small Spacecraft Constellation Concept for Mars Atmospheric Radio Occultations***

**S.W. Asmar, A.J. Mannucci, C.O. Ao, M.M. Kobayashi**

**J. Lazio, A.D. Marinan, R.A. Preston, W. Williamson**

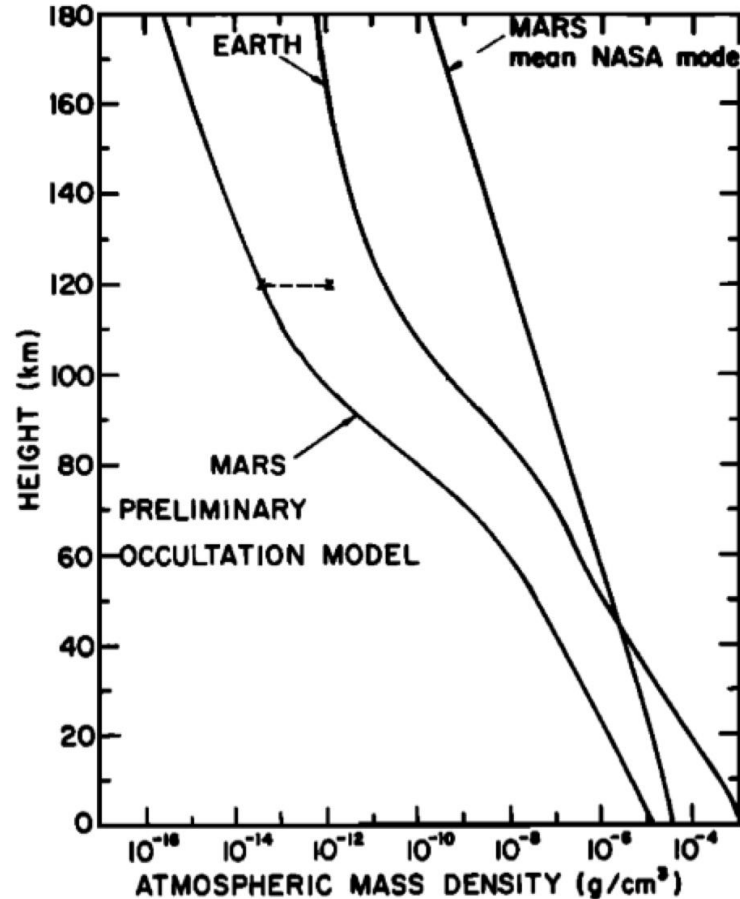


**Jet Propulsion Laboratory**  
California Institute of Technology

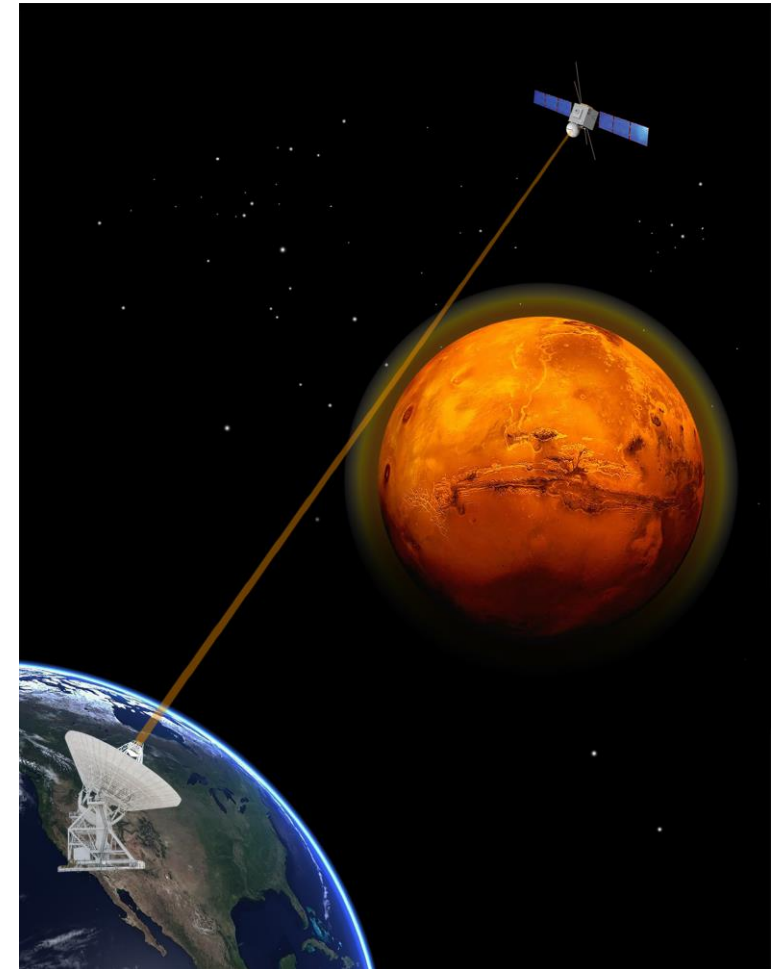
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Government sponsorship acknowledged.

# Atmospheric Radio Occultations

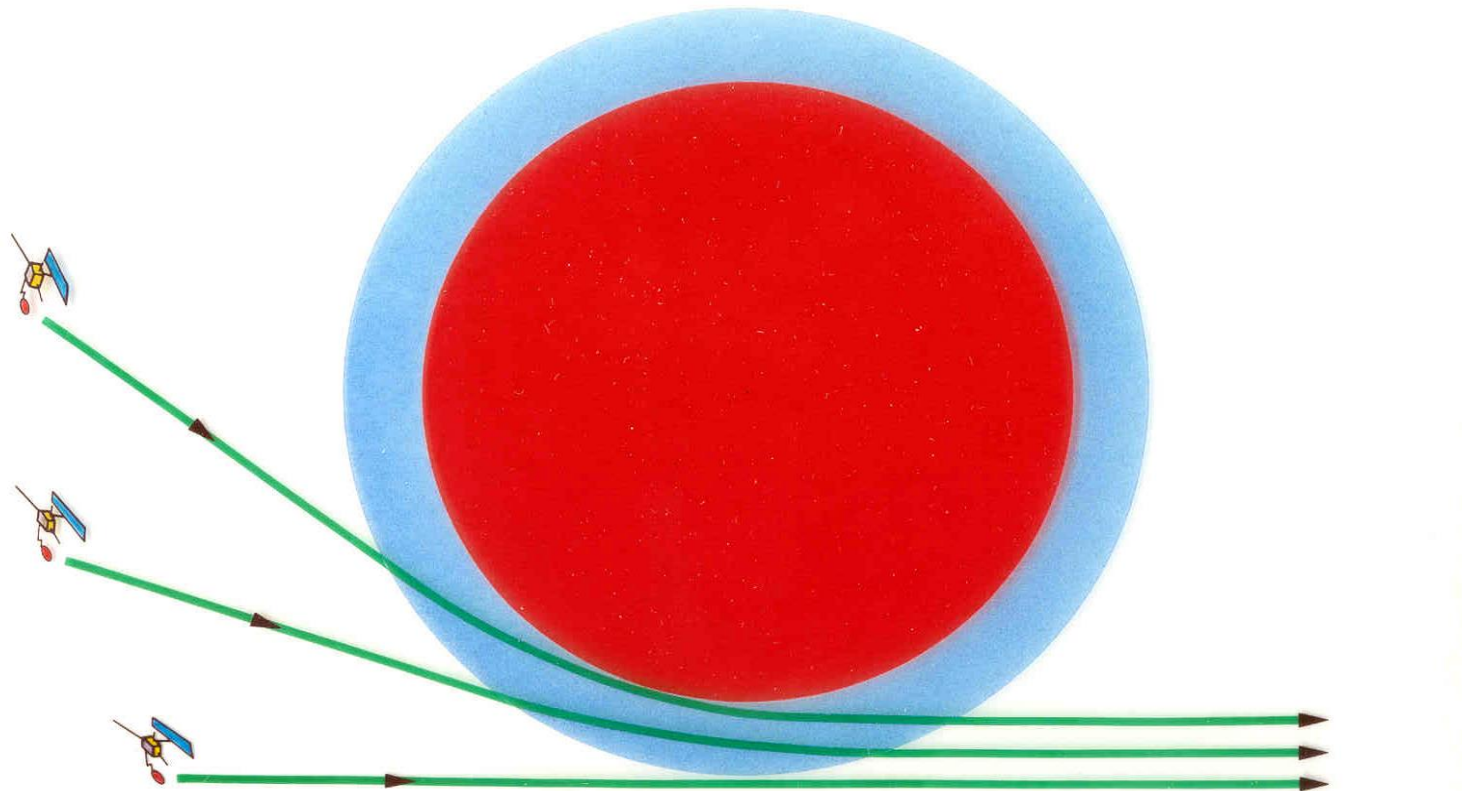
- Temperature-pressure profiles
- Ionospheric density and structure
- First results in 1965
- Technique later applied to occultations of planetary rings, plasma torus, and solar corona



Kliore et al, 1965

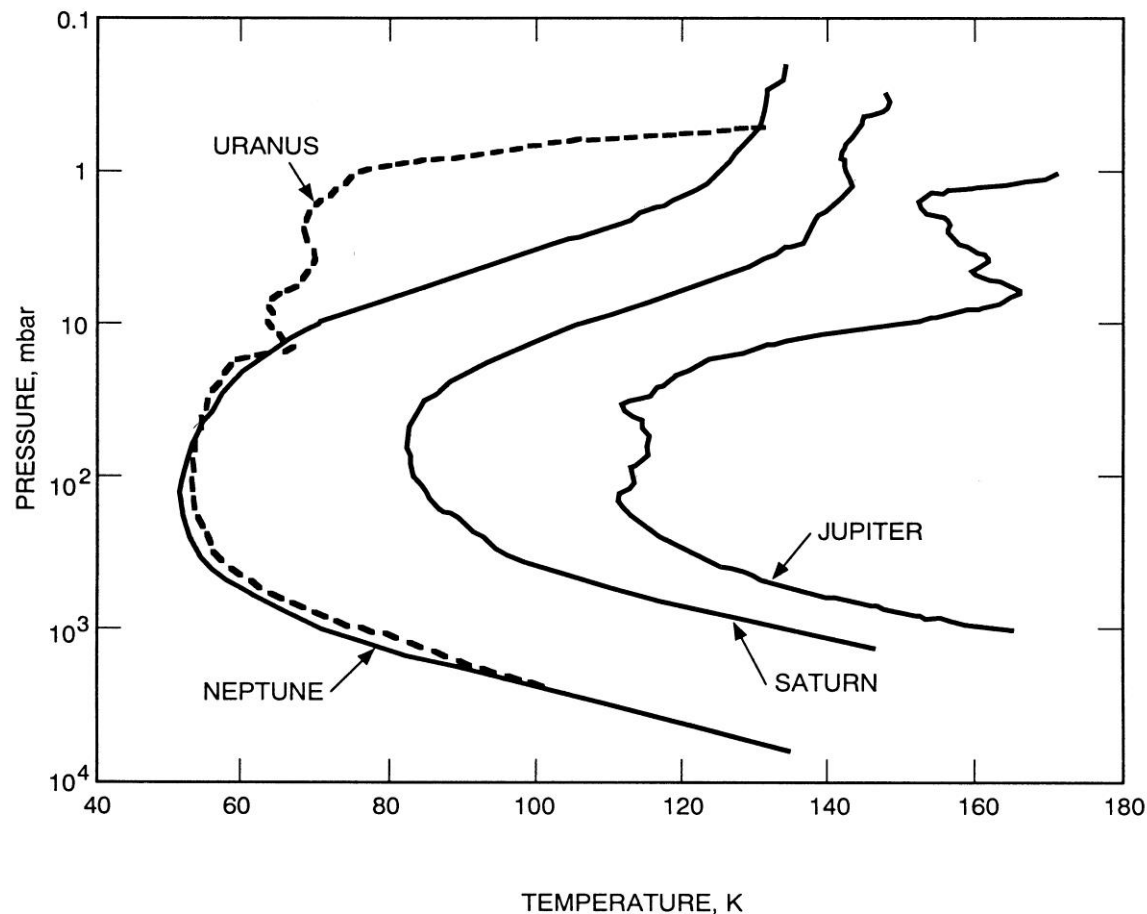


Phase Change → Path Length → Refractive Angle → Refractivity  
→ Number Density → Column Pressure → Temperature Profiles

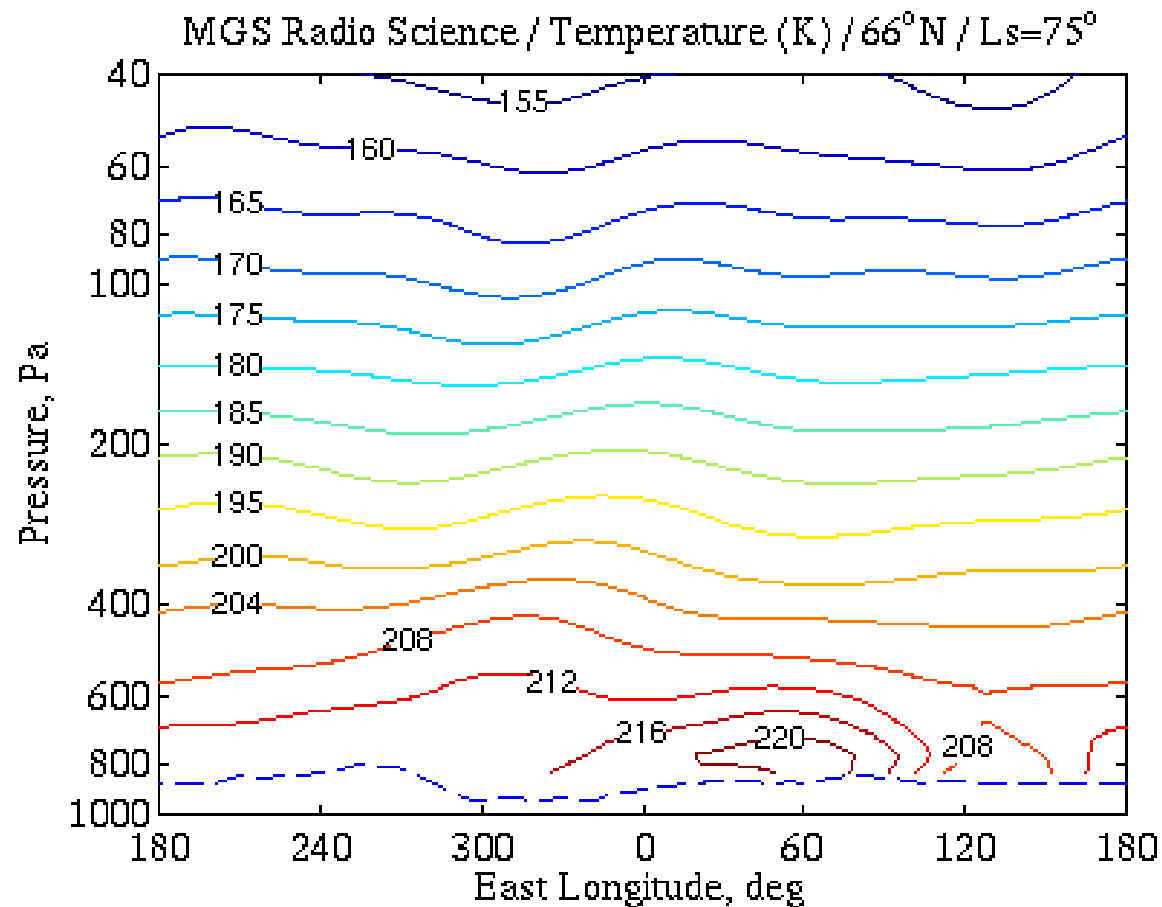


Source: M. Patzold

# Flybys at Outer Planets vs. a Decade in Mars Orbit

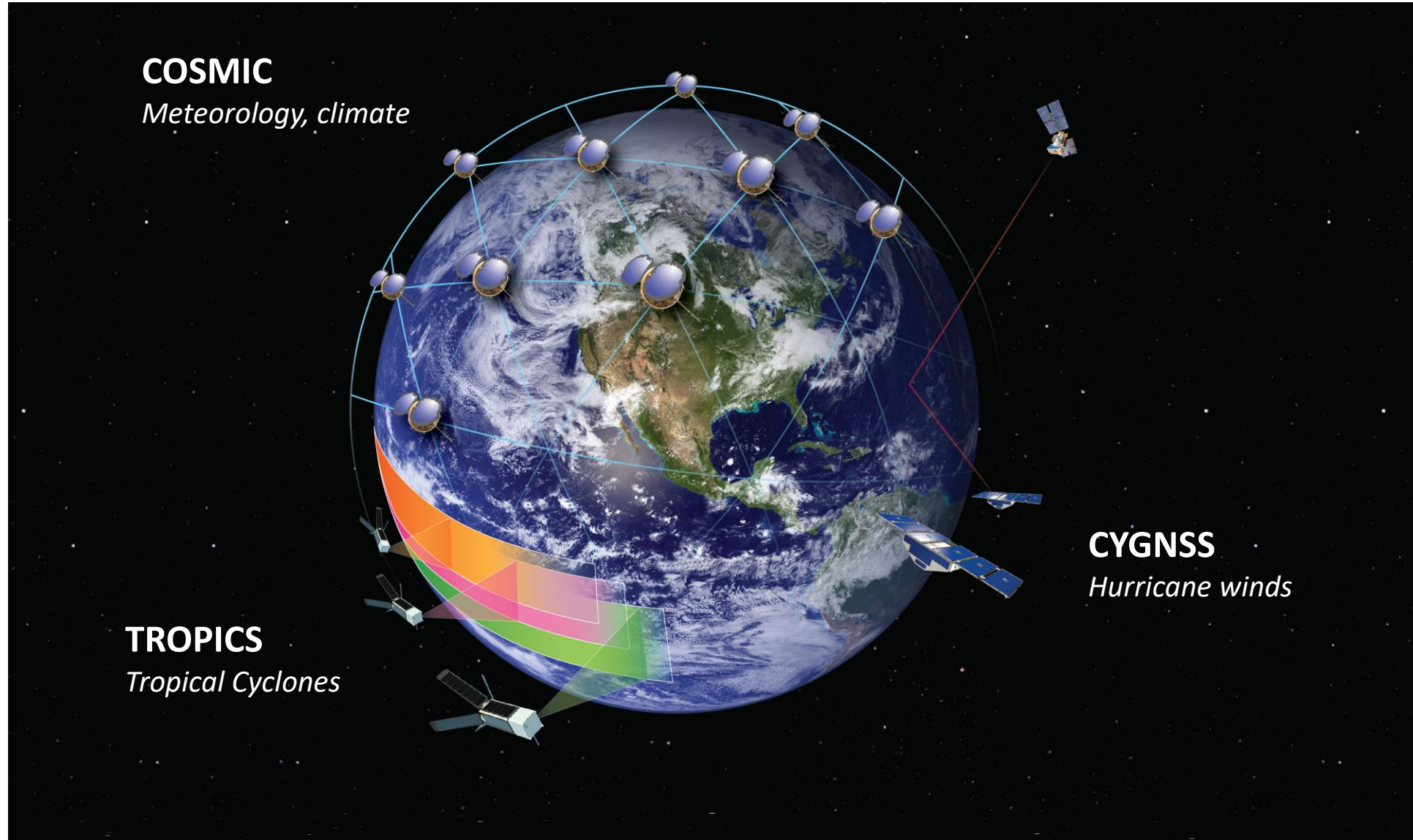


Temperature profiles for the giant planets derived from radio occultation data acquired with the Voyager spacecraft (from Lindal, 1992)



Source: D. Hinson

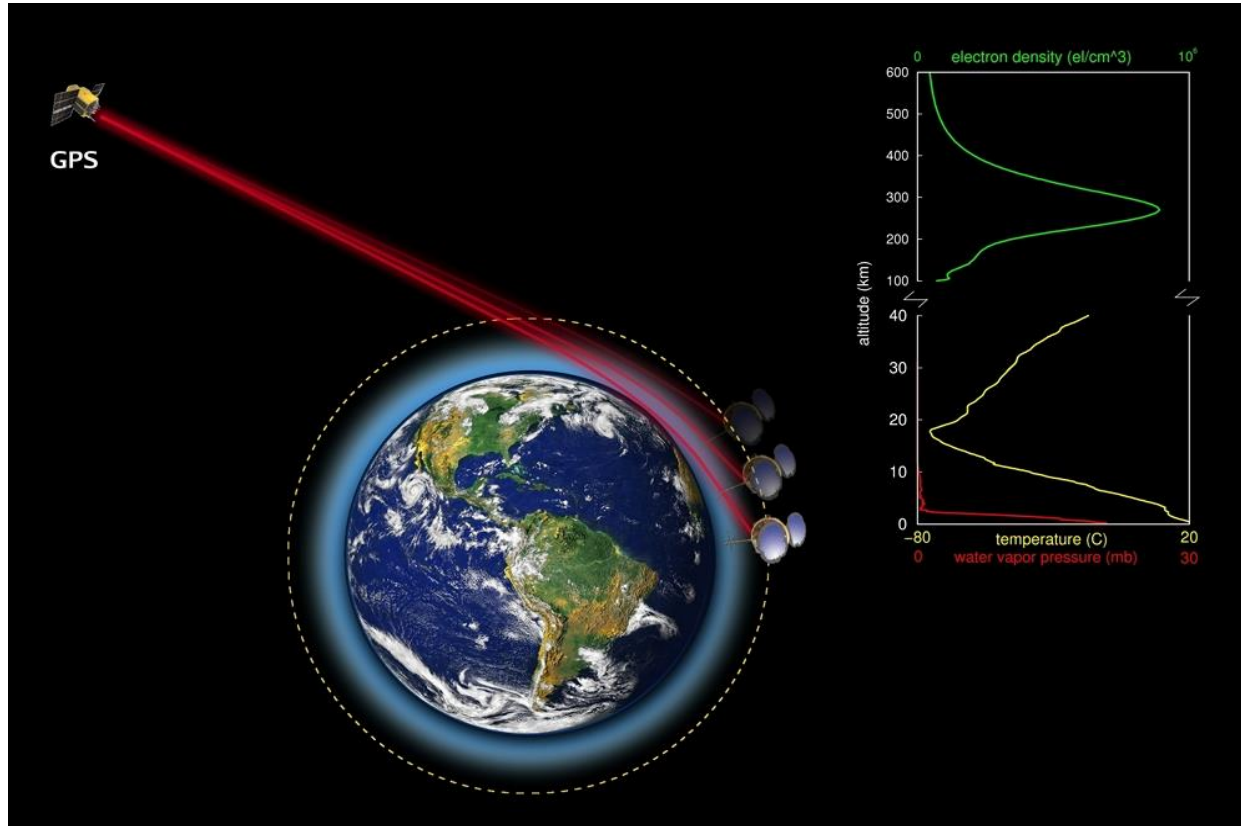
# Earth Atmospheric Occultations via Constellations





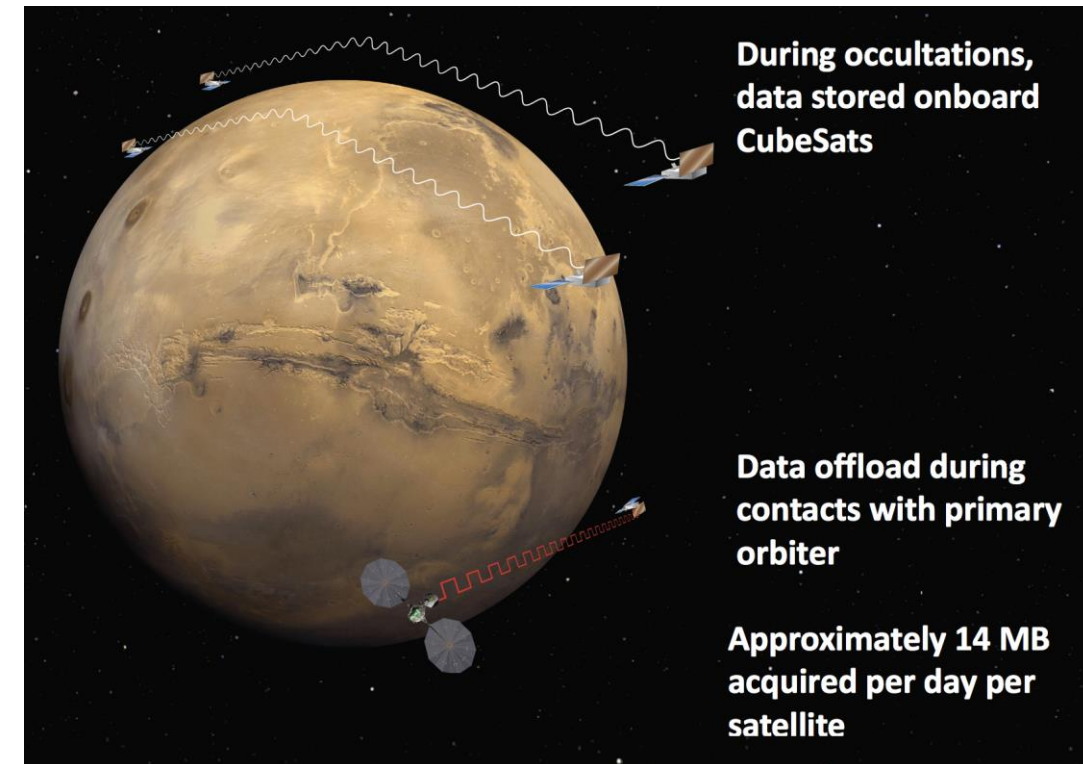
# Earth Atmospheric Occultations Example Results

- Anthes et al., 2008, COSMIC/FORMOSAT-3
- 6 microsattellites launched into 512-km orbit providing 1500 RO daily
- Global 3-D coverage: 40 km to surface
- 70% - 90% of RO reach to ~1 km of surface
- Vertical resolution: ~100 m in lower troposphere
- Independent height, pressure, and temperature data
- High accuracy: averaged profiles to < 0.1 K



# From *Spacecraft-to-DSN* to *Spacecraft-to-Spacecraft*

- 5 decades of planetary spacecraft links to Earth
  - *Reversed DSN-to-spacecraft for NH at Pluto*
- Earth science community advanced technique via GPS satellites transmitting to science spacecraft
- Carry *crosslink* concepts to planetary atmospheres for tremendous advantages
- Science motivation & assist human missions (✓)
- Need mission design & navigation (✓)
- Need radio instrumentation (✓)
- Need small spacecraft platform (✓)
- Need a demo of concept (✓)
- Ready for science mission implementation

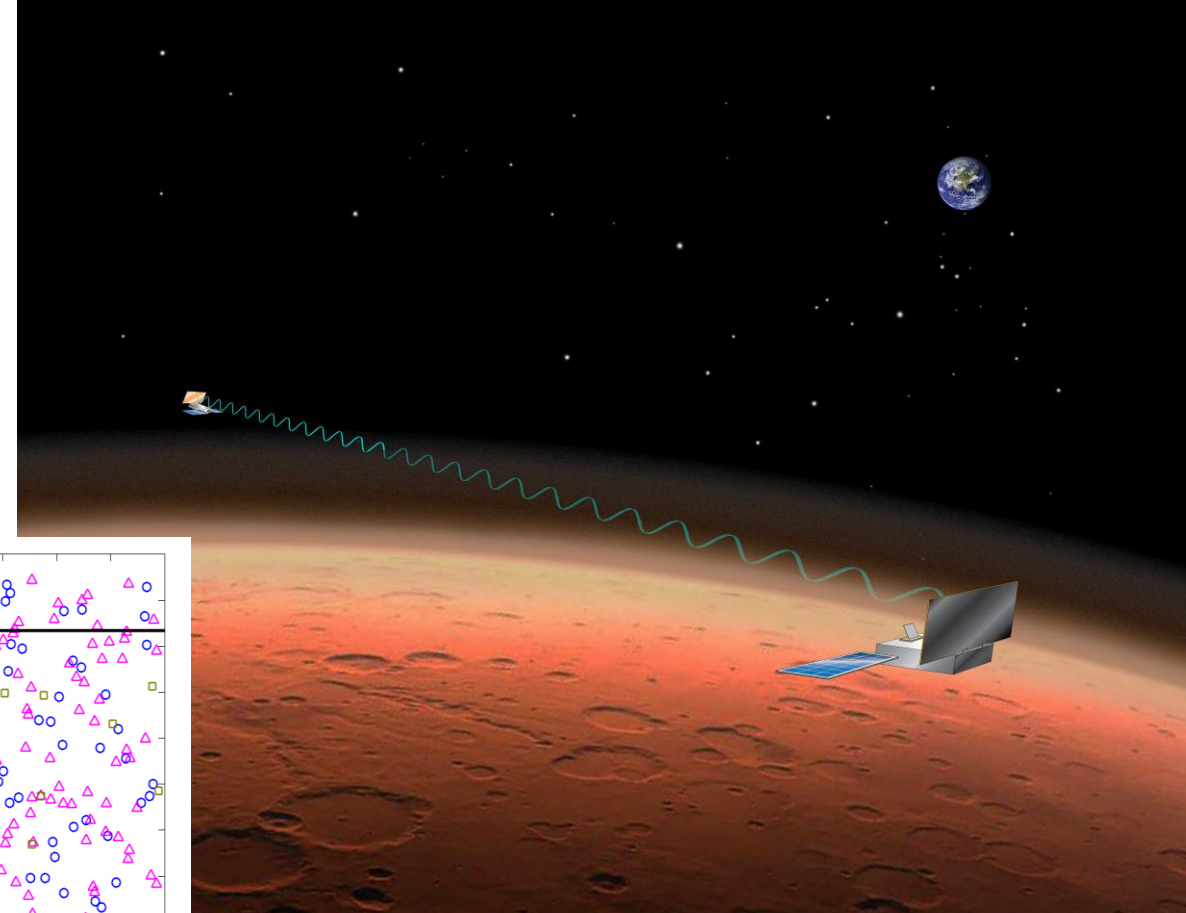
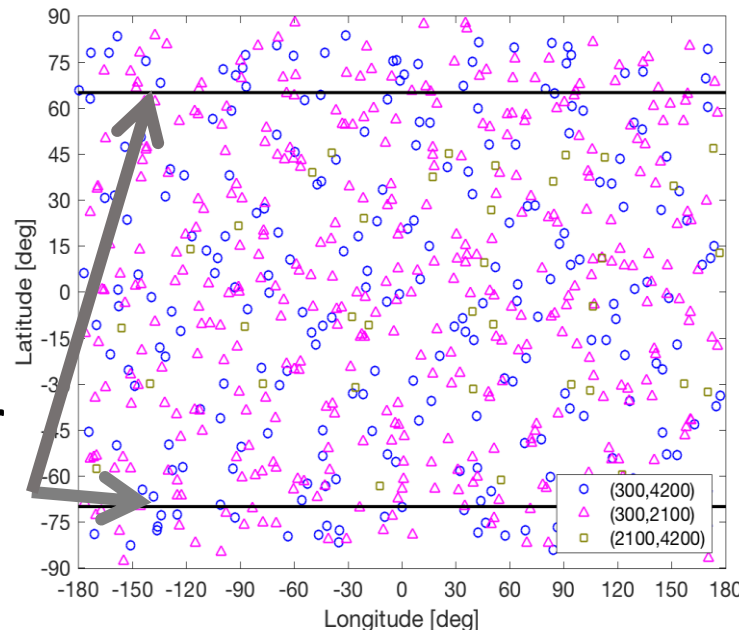


*Pre-Decisional Information -- For  
Planning and Discussion Purposes Only*

# Advantages of Crosslink Radio Occultation

- Improve temporal & spatial resolution
- Rapid global coverage with diurnal, seasonal, annual repeatability
- High vertical resolution of T-P profiles
- Not obscured by dust & profiles reach surface
- Possibly higher SNR

- Three CubeSats at 3 altitudes
- Pairwise occultation locations
- One week acquisition time
- Compare to single polar orbiter one week spacecraft-to-DSN covering polar regions only



*Pre-Decisional Information -- For Planning and Discussion Purposes Only*



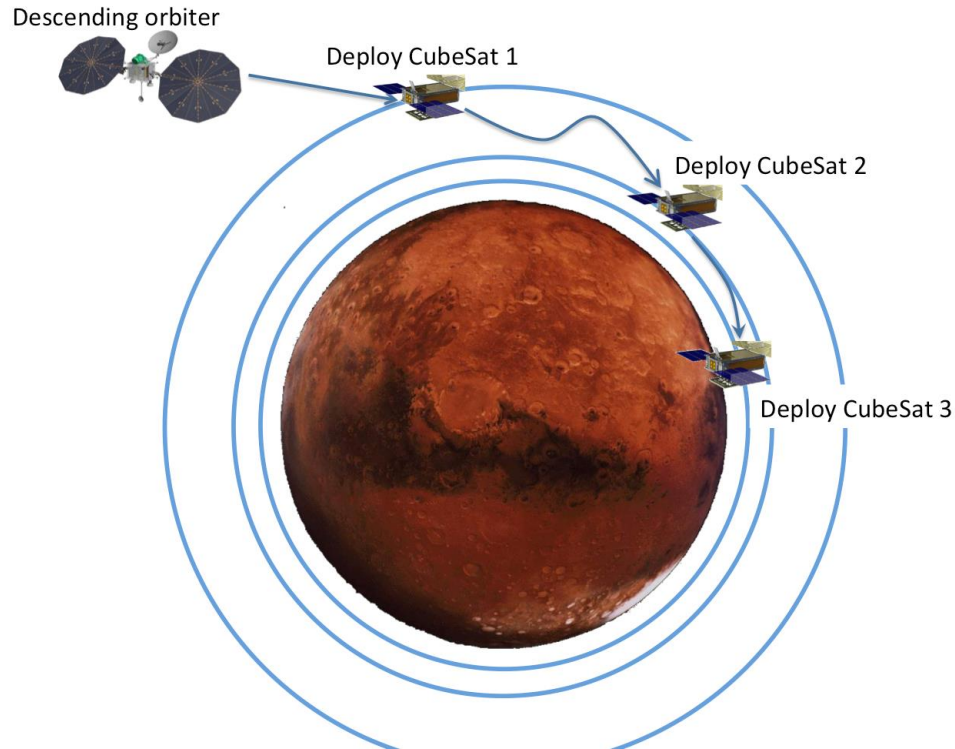
# MEPAG Goal II: Understand Processes & History of Climate

Objectives	Sub-objectives
<b>A.</b> Characterize the state of the present climate of Mars' atmosphere and surrounding plasma environment, and the underlying processes, under the current orbital configuration.	<b>A1.</b> Constrain the processes that control the present distributions of dust, water, and carbon dioxide in the lower atmosphere, at daily, seasonal and multi-annual timescales.
	<b><u>Investigation A1.1:</u></b> Measure the state and variability of the <b>lower atmosphere</b> from turbulent scales to global scales (High Priority).

## MEPAG Goal IV: Prepare for Human Exploration

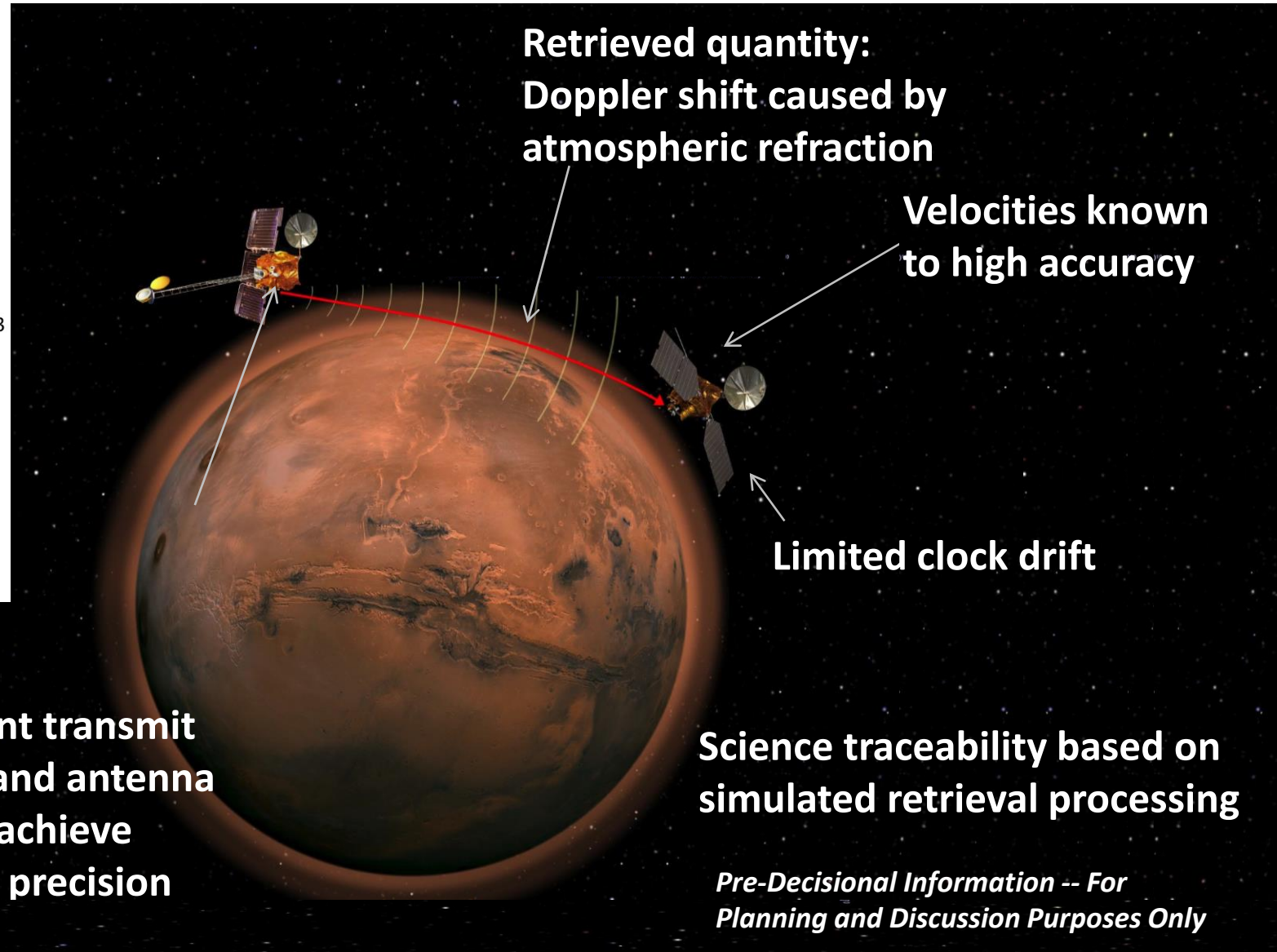
<b>B.</b> Obtain knowledge of Mars sufficient to design and implement a human mission <b>to the Martian surface</b> with acceptable cost, risk, and performance.	<b>B1.</b> Determine the aspects of the atmospheric state that affect Entry, Descent, & Landing (EDL) design, or atmospheric electricity that may pose a risk to ascent vehicles, ground systems, and human explorers.
	<b><u>Investigation B1.2:</u></b> Monitor surface pressure and near surface meteorology over various temporal scales (diurnal, seasonal, annual), and if possible in more than one locale (High Priority).

# Measurement Physics Drives the Design Concept



- Deploy during descent of delivery orbiter
- Differential precession of orbit planes achieves desired variety of orbits
- Candidate altitudes: 4200, 2100, & 300 km

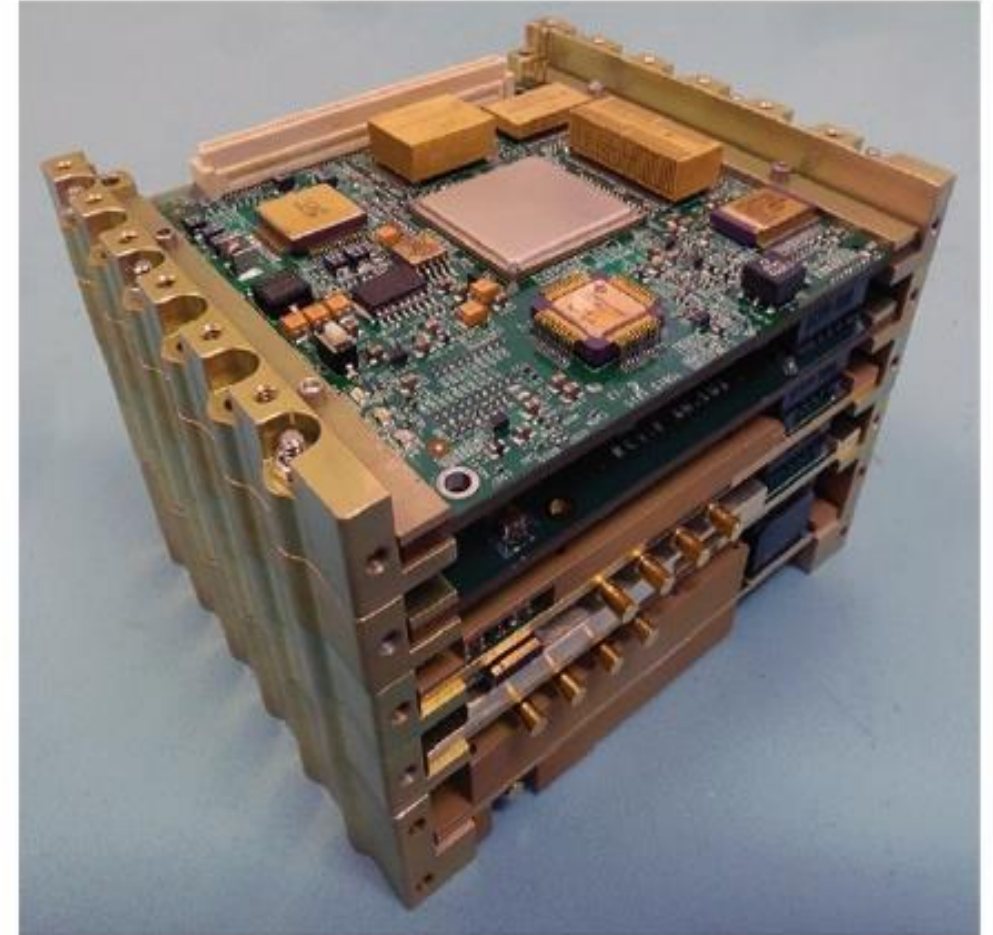
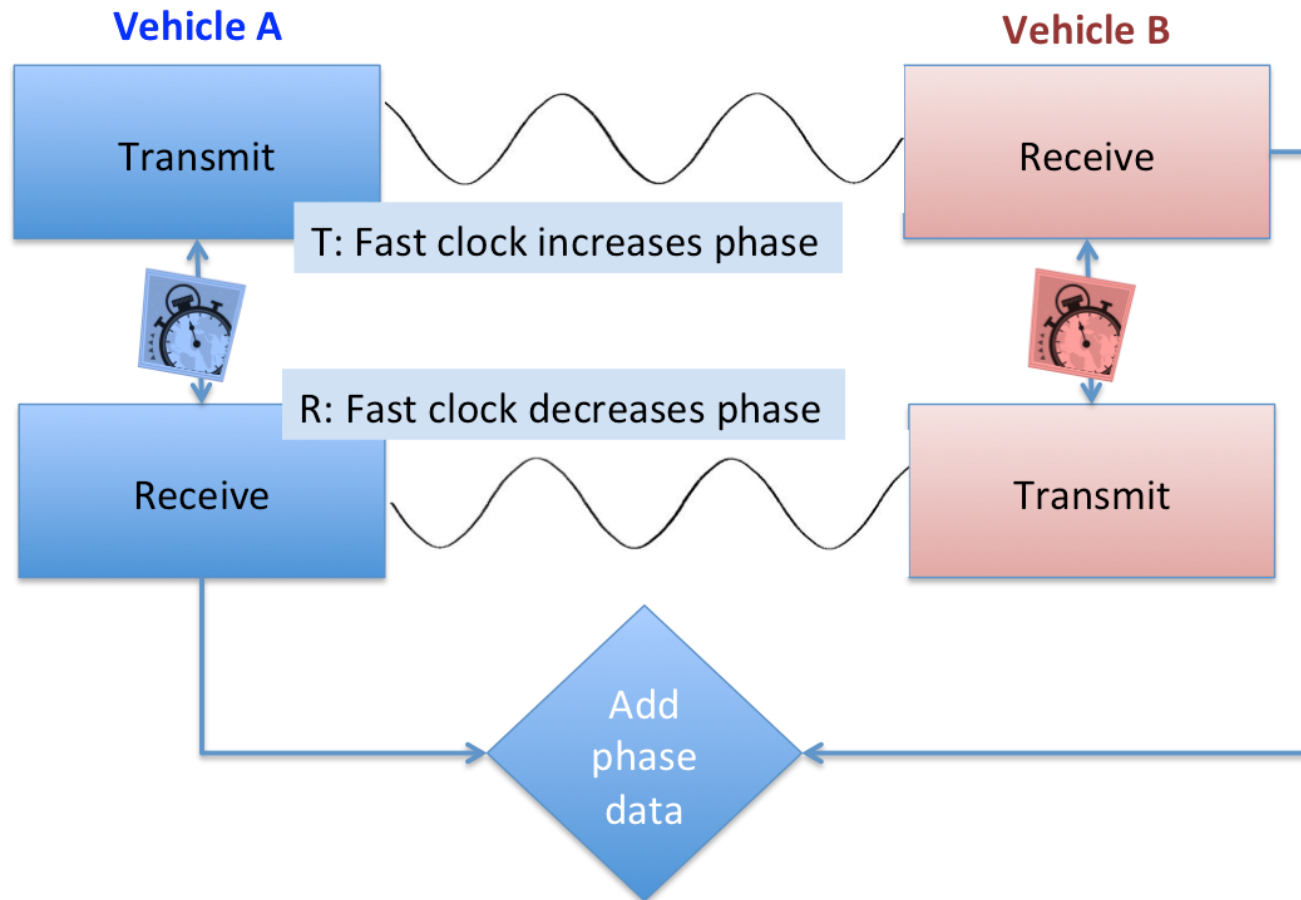
**Sufficient transmit power and antenna gain to achieve needed precision**



**Science traceability based on simulated retrieval processing**

*Pre-Decisional Information -- For Planning and Discussion Purposes Only*

# Iris Radio with Dual-One Way Range Rate

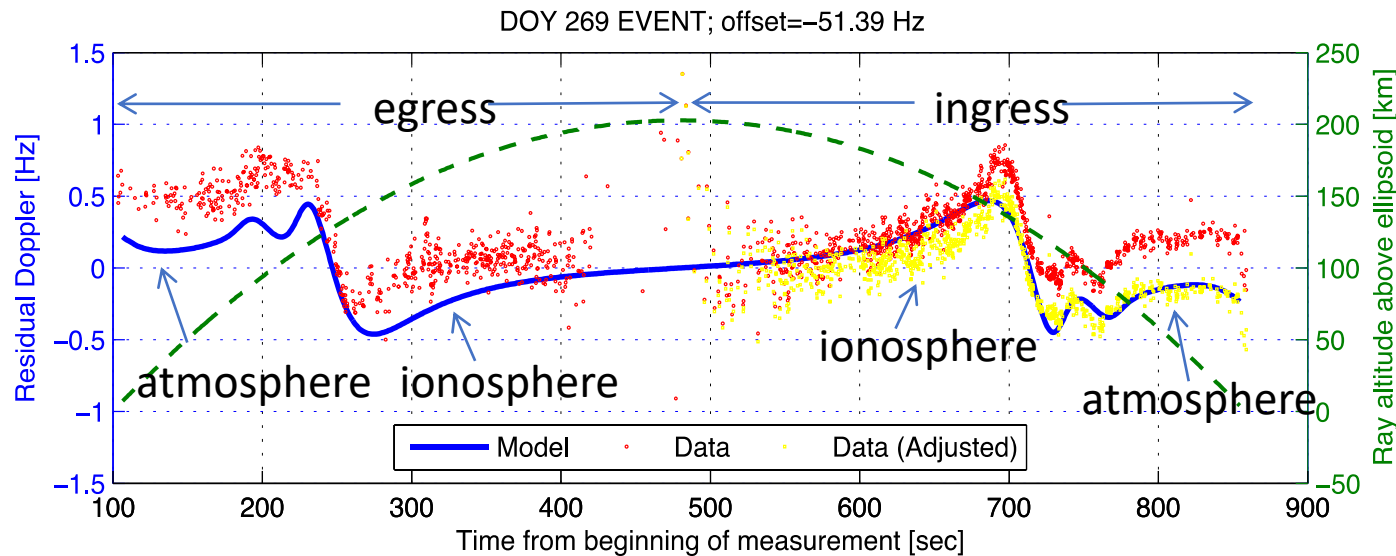


**Simultaneous X-band transmit/receive capability  
Navigation requirements achieved with crosslink  
measurements above atmosphere and links to  
orbiter**



# Demonstration: UHF Link from Odyssey to MRO

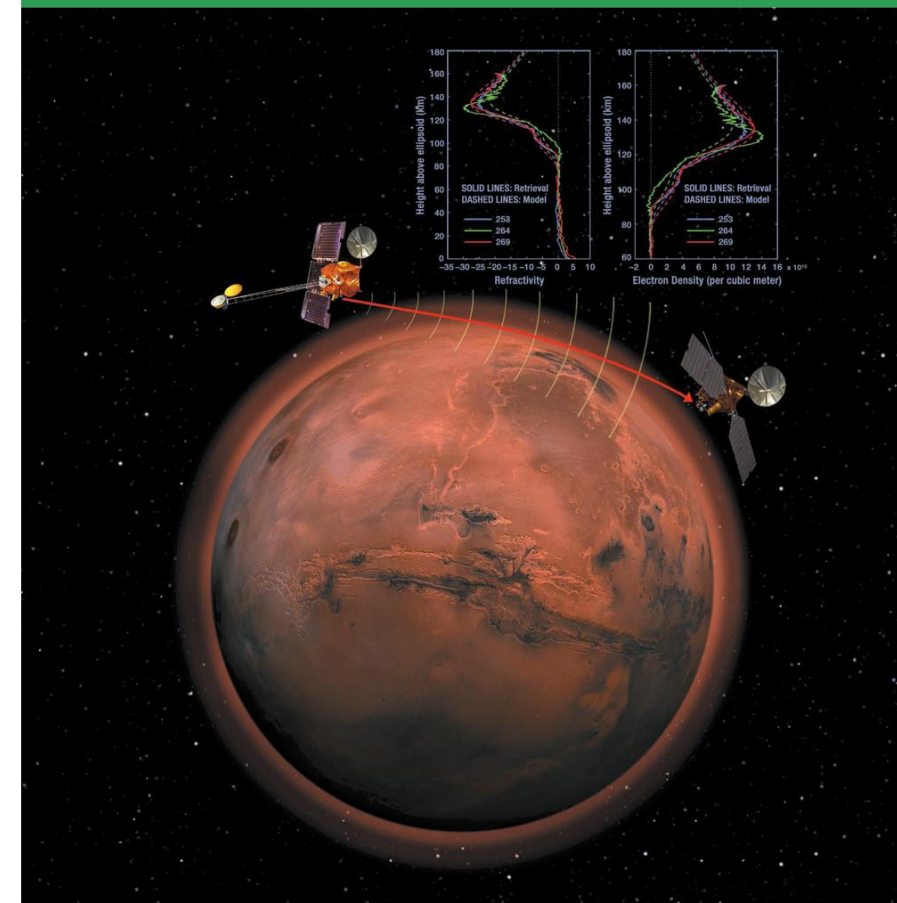
- Found geometries where orbiters communicating with rover on surface were in line-of-sight for a crosslink ( $\lambda \sim 75$  cm); made possible due to flexible design of the Electra software-defined radio.



## Radio Science

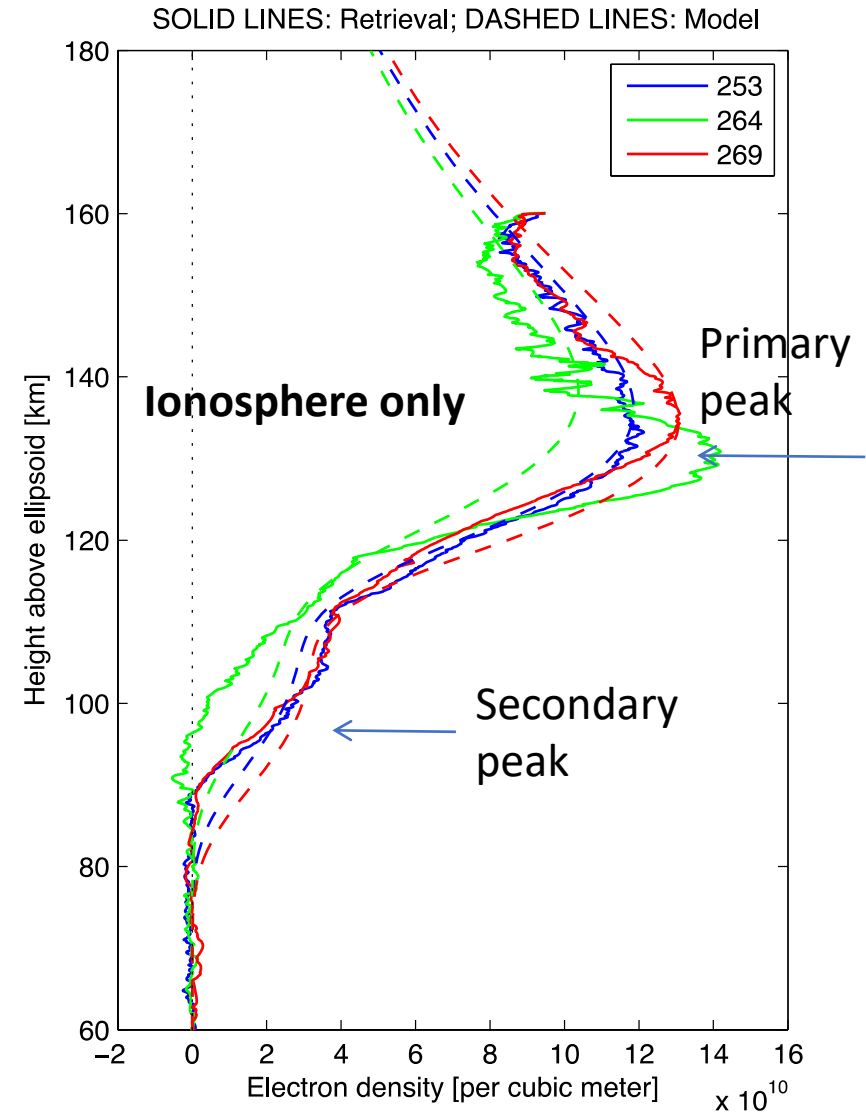
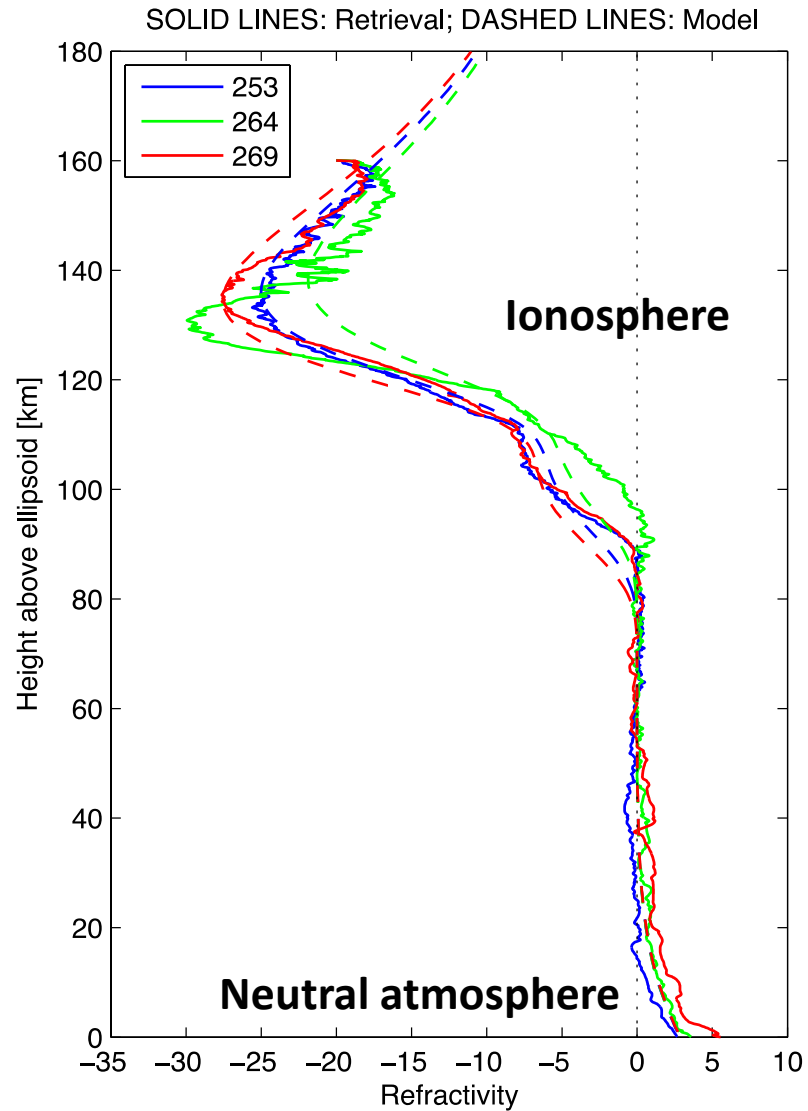
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# Ionospheric Retrievals from UHF Radio



# Start with the *Mars Cube One* (MarCO) Design



- 6U, soon to be first planetary CubeSat
- Make radio system the science instrument
- Cost savings, lower risk
- MarCO's X-band reflect-array (relay of 8 kbps) replaced by patch antenna
- Other simplifications

